

## Two Dimensional Turbulence and Typhoon Concentric Eyewall Formation

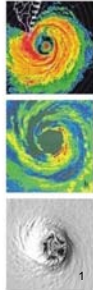
郭鴻基

教育部國家講座教授  
臺大終身職特聘教授

6/24/2009 中研院環變中心



The two-dimensional turbulence, which may be stemmed either from the fluid rotation or the fluid internal stratification, has been a paradigm for geophysical fluid dynamics. It is a remarkable fact that for any type of random initial state or external forcing, a two-dimensional fluid will rapidly organize itself into a system of coherent, interacting vortices swimming through a sea of passive filamentary structure produced from earlier vortex interactions. This discipline has also been instrumental in the development of single charge plasma physics, galaxy spiral, the Great Red spot, the ozone-hole problem and the typhoon dynamics. A brief review of 2D and 3D turbulence will be given. The typhoon merger dynamics, the eyewall rotation and mixing, the successive formation of tropical vortices and the concentric eyewall formation dynamics will be presented. The essence of the scientific computing in 2-D turbulence will be addressed. An efficient domain decomposed Chebyshev spectral method will be discussed.



一杯咖啡，古今往事盡付笑談中。

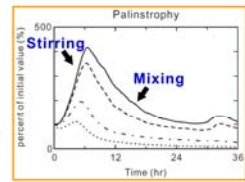
The best part of waking up, is the vortex in your cup!

$$\frac{D\theta}{Dt} = \frac{\partial\theta}{\partial t} + \vec{v} \cdot \nabla\theta = v\nabla^2\theta$$

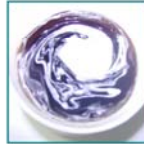
$$C = \frac{1}{2} \int \nabla\theta \cdot \nabla\theta dV$$

$$\frac{dC}{dt} = \int (\vec{v} \cdot \nabla\theta) \nabla^2\theta dV - \nu \int (\nabla^2\theta)^2 dV$$

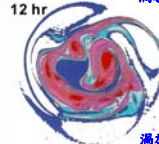
Stirring Mixing



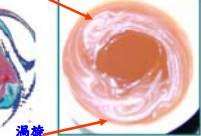
Coffee with white



12 hr

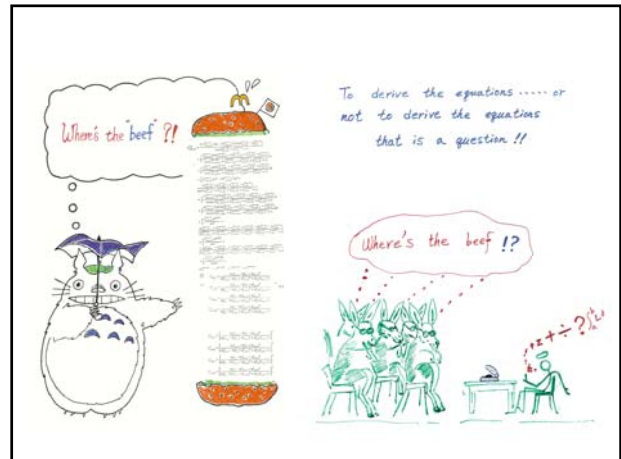


渦旋

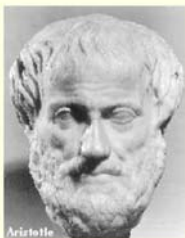


渦旋

Politics are for the moment  
An equation is for eternity



## Aristotle's Meteorologia



Aristotle (384-322 BC) was a past master at asking questions.

He wrote the first book on Meteorology, the *Μετεωρολογία* (μετεωροι: Something in the air)

This work dealt with the causes of various weather phenomena and with the origin of comets.

While a masterly speculator, Aristotle was a poor observer: for example, he believed that the lightning followed the thunder!

形而上學：憑藉第一原因，一切事物方能知曉，但其本身是自明的。

## Galileo's Thermometer



Heat and Buoyancy

溫度與浮力

Potential temperature


The Galileo Thermometer is a popular modern collectable and an attractive decoration.

As temperature rises, the fluid expands and its density decreases.

The reduced buoyancy causes the glass baubles to sink, indicating temperature changes.

$$\rho' = -\alpha T' \quad \alpha = -\frac{\partial\rho}{\partial T} > 0 \text{ in general}$$

Peter Lynch




**Isaac Newton**

**Principia 1687**

**Nature and nature's law  
lay hid in night,  
God said,  
Let Newton be,  
and all was light. A. Pope**


7

**Edmund Halley (1656-1742)**



Edmund Halley was a contemporary and friend of Isaac Newton.  
He was largely responsible for persuading Newton to publish his *Principia Mathematica*.

**Halley and his Comet**



Halley's analysis of what is now called Halley's comet is an excellent example of the scientific method in action.


**A Tricky Question??**

If the astronomers can make Accurate 76-year forecasts, Why can't the meteorologists do the same?

Size of the problem  
大氣海洋自由度無限 + 熱力學

Order versus chaos  
大氣海洋的混沌、蝴蝶效應

8



**Halley 1686**

First proposed the atmospheric motion is connected with the **distribution of sun heat** (follow the sun in the daily scale; thus wind is westward.)

**Halley 1686**

**Sun heat**

**夸父逐日  
東風逐日**

環流在東西方向均勻

航海時不同經度判定


**Y-Z Profile**

9

**Torricelli's Barometer**

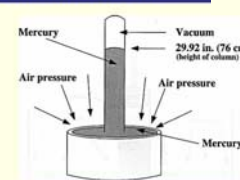
**Galileo's "suction vacuum"**  
**Torricelli's air pressure "pushing"**

**Air weigh 1000kg/m<sup>2</sup>**



Evangelista Torricelli (1608-1647), a student of Galileo, devised the first accurate **barometer**.

**Barometric Pressure**



The relationship between the height of the mercury column and the character of the weather was soon noticed.

10

$P = \text{force/area} = ML^{-1}M^{-2}L^{-2}$   
 $= ML^{-1}M^{-2}$   
 $1N/m^2 = 1 \text{ Pascal}$   
 $1 \text{ millibar} = 100 \text{ Pa} = 1 \text{ hPa}$

**Material volume; Fluid parcel**

Euler and Bernoulli regarded the **pressure as a field acting upon an imaginary closed boundary** and mechanically equivalent to the action of the fluid exterior to the boundary upon that interior to it.

Previous scientists had regarded the pressure as the weight of a unit column of fluid.





Figure 21 A cartoon by Sidney Harris, a noted chronicler of humor in science. ("Hydrodynamics" Bernoulli is our man.)

11



**Euler's Equations 18世紀**

$$D/Dt = \frac{\partial \vec{v}}{\partial t} + \vec{v} \cdot \nabla \vec{v} = -\frac{1}{\rho} \nabla p - \nabla \Phi$$

Momentum equation

$$\frac{d\rho}{dt} + \rho \nabla \cdot \vec{v} = 0$$

Continuity equation

4 equations 5 unknowns

$$p = p(\rho)$$


Constitutive equation  
Equation of State

**流體力學之父**

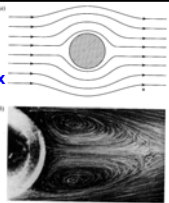
**流體塊  
壓力梯度力  
座標轉換  
質量守恒**

**Nonlinear Partial Differential Equations**  
非線性偏微分方程式 PDE

12



**D'Alembert Paradox**




1717-1783  
D'Alembert Wave Solution

$$\frac{\partial^2 u}{\partial t^2} = c^2 \frac{\partial^2 u}{\partial x^2}$$

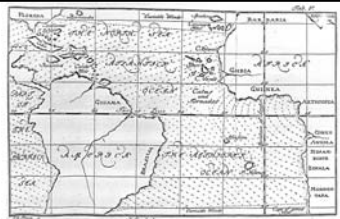
$$u(x,0) = f(x), u_x(x,0) = g(x)$$

$$u(x,t) = \frac{1}{2} [f(x-ct) + f(x+ct)] + \frac{1}{2c} \int_{x-ct}^{x+ct} g(\xi) d\xi$$

13



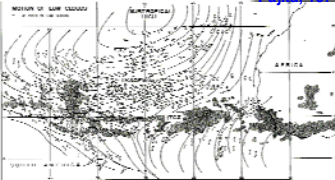
1746



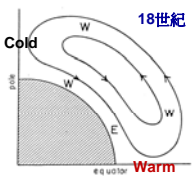
D'Alembert 1746

**Solar and Lunar Force**

Math. Model for Atmospheric Motion in aqua-planet (Won the 1746 Berlin Academy's Award; Euler's endorsement)



Fujita, 1971



18世紀

Cold

Warm

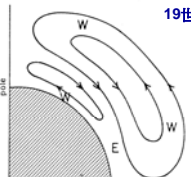
Fig. 2. A schematic view of a meridional cross-section of the general circulation as visualized by Hadley (1735). Streamlines indicate the meridional and vertical flow, while letters E and W indicate regions of easterly and westerly flow.

Hadley (1685-1758)

Distribution of sun heating (north and south; seasonal scale)

Earth rotation (conservation of angular momentum) 角動量守恆

Energy transport and balance



19世紀

W

E

Fig. 3. The same as Fig. 2, but for the general circulation as visualized by Thomson (1857) and Ferrel (1859).

Ferrel (1859), Thomson (1857)


Centrifugal force 離心力

Thermal wind balance 熱力風平衡

Surface friction

Coriolis 1835

15



**Pierre-Simon Laplace**  
1740-1827

Observational net work 1800-1815  
氣象觀測網

Tidal equations

Laplacian  $\nabla^2 \psi = f$

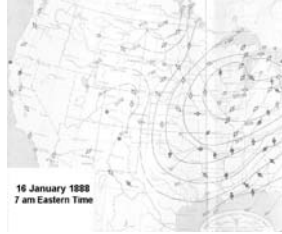
Hydrostatic approximation  $\frac{\partial p}{\partial z} = -\rho g$

Adiabatic sound speed  $\left(\frac{\partial p}{\partial \rho}\right)_\theta = c_s^2$   
絕熱聲速

Laplace Transform

Astronomer

Black hole and gravitational collapse

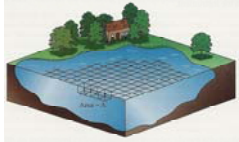


16 January 1888  
7 am Eastern Time

16

## Estimate Avogadro's Number


Benjamin Franklin (1773)



Oil spreads on water  
→ molecular size  
→ Avogadro's number

- Molecular size  
 $l = \frac{V}{A} = \frac{4.9 \text{ cm}^3}{2.0 \times 10^7 \text{ cm}^2} = 2.4 \times 10^{-7} \text{ cm}$
- Number of molecules  
 $N = \frac{A}{l^2} = \frac{2.0 \times 10^7 \text{ cm}^2}{(2.4 \times 10^{-7} \text{ cm})^2} = 3.5 \times 10^{20} \text{ molecules}$
- Mass of the oil  
 $m = V \times D = 4.9 \text{ cm}^3 \times 0.95 \frac{\text{g}}{\text{cm}^3} = 4.7 \text{ g}$
- Number of moles of oil  
Moles of oil =  $\frac{4.7 \text{ g}}{200 \text{ g/mol}} = 0.024 \text{ mol}$
- Avogadro's number  
Avogadro's number =  $\frac{3.5 \times 10^{20} \text{ molecules}}{0.024 \text{ mol}} = 1.5 \times 10^{22}$

Now we know:  $N_A = 6.022142 \times 10^{23} \text{ /mol}$



**Ideal Gas Law Equation of State 理想氣體方程**

- 1662, Boyle law,  $PV = c$  when  $T = c$ .
- 1787, Charles law,  $V/T = c$  when  $P = c$ .
- 1803, Gay-Lussac law,  $P/T = c$  when  $V = c$ .
- 1811, Avagadro, 1 mole gas is 22.4 l in volume.


Universal Gas Constant  
 $R^* = 8314.3 \text{ J/(deg} \cdot \text{kmol)}$

$PV = n R^* T$   
 $PV = m/M R^* T$   $P = m/V R^*/M T$   
 $P = \rho R T, R = R^*/M$   
 $R_u = 287 \text{ J/deg} \cdot \text{kg}$  ( $R^*/M_u$ )  
 $R_v = 461 \text{ J/deg} \cdot \text{kg}$  ( $R^*/M_v$ )


$P = f(V, T)$   
 $V = h(P, T)$   
溫度可以改變壓力或體積，熱可以做功

18

Development of Thermodynamics 熱力學 雲微物理  
 19 century  
**第一定律 能量作功，能量守恆**  
 First law: Energy is what makes it go and energy is conserved.  
 $\Delta Q = \Delta U + \text{WORK}$   
 Second law: Entropy tells it where to go!  
**第二定律 時間之矢，自然單向**  
 Joule, Rudolf Clausius, Lord Kelvin and others  
**宏觀 微觀**  
 Macro --- Micro  
 Classical and Statistical Thermodynamics 統計熱力學  
 Ludwig Boltzmann, 1844-1906, whose work led to an understanding of the macroscopic world on the basis of molecular dynamics.  
S = k Log W



**Fourier 1768-1830**  
 1824 Why the earth not heating up when receive sun energy continuously?  
**Heat emission or diffusion (by IR)**  
 His calculations showed a very cold surface (No green house effect)



**Tyndall 1864**  
 H<sub>2</sub>O is a green house gas

**Arrhenius 1896**  
 CO<sub>2</sub> green house effect, but were dismissed by scientists [WHY??]

**Lord Rayleigh 1881**  
 Rayleigh scattering  
**Why the sky is blue??**

**Chandrasekhara Raman 1928**  
 Raman scattering  
**Color of the sea**


**Jame Chappuis 1881**  
 Chappuis band  
 dismissed by scientist

**Edward Olson Hulburt 1956**  
 Twilight blue  
 O<sub>3</sub> layer

Fovell, Taipei, 2008

20

The profound study of nature is the most fertile source of mathematical discoveries.



Fourier, Jean Baptiste Joseph  
 1768-1830

1807 at age 39; argued with Lagrange and Laplace on the representation of a triangle wave with cosine and sine functions.

$$f(x) = \sum f_k e^{ikx}$$

$$\hat{f}_k = \frac{1}{2\pi} \int_0^{2\pi} f(x) e^{-ikx} dx$$

$$f(x) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} \hat{f}_k e^{ikx} dx$$

$$\hat{f}_k = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} f(x) e^{-ikx} dx$$

Heat emission or diffusion (by IR)  
 His calculations showed a very cold surface (No green house effect)

**f(x) does not have to be analytical; f(x) does not have to be periodic.<sup>21</sup>**

**Planck, Unwilling Revolutionary: the idea of quantization**  
 1900  
**Hall of Fame in Science**  
**Gravitational Law**  
**Blackbody Radiation**  
**E= MC^2**

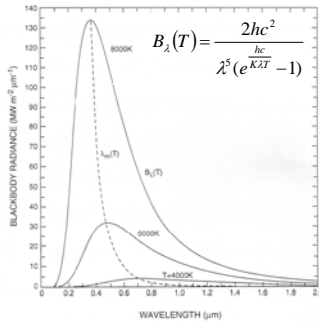




Figure 8.7 Spectra of emitted intensity  $B_\lambda(T)$  for blackbodies at several temperatures, with wavelength of maximum emission  $\lambda_m(T)$  indicated.



1858-1947

**Multiple Scale Interactions**  
**多重尺度交互作用**  
**Current vortices**

**Leonardo da Vinci 1452-1519**  
 Observe the movement of the surface of the water which resembles that of hair which has **two motions**, of which one depends on the weight of the hair and the other on the direction of the **curls**. Thus water forms eddying whirlpools of which one part depends on the predominant **current** and the other on the incidental motion and the return flow."



**Brief History of Fluid Dynamics**

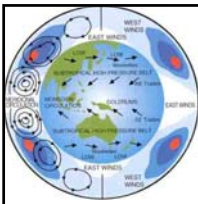
<b>Newton</b>	<b>1700s</b>	Viscosity Law of motion for a particle	18世紀
<b>Euler</b>	<b>1750s</b>	Equations for inviscid flow, Law of motion applied to fluids	
<b>D. Bernoulli</b>			19世紀
<b>Navier</b>	<b>1827</b>	Equations for <b>viscous</b> fluid flow	
<b>Stokes</b>	<b>1845</b>		
<b>Boussinesq</b>	<b>1877</b>	Turbulent mixing, <b>eddy viscosity</b>	20世紀
<b>Reynolds</b>	<b>1880</b>	Transition to turbulence, Reynolds Number, <b>Reynolds stress</b>	
<b>G. I. Taylor</b>	<b>1915-1970</b>	Geophysical flows, rotating flows	
<b>Prandtl</b>	<b>1904</b>	<b>Boundary Layer</b>	

$$\tau_{ij} = -\rho_0 \overline{u_i' u_j'}$$

**Reynolds stress**

$$\frac{\partial \bar{u}_i}{\partial t} + \bar{u}_j \frac{\partial \bar{u}_i}{\partial x_j} = -\frac{1}{\rho_0} \frac{\partial \bar{p}}{\partial x_i} + \frac{g_i}{\theta_0} \delta_{i3} - \frac{\partial}{\partial x_j} (\overline{u_i' u_j'}) + \nu \frac{\partial^2 \bar{u}_i}{\partial x_j^2}$$

24

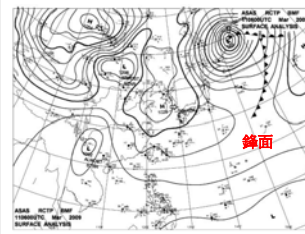


20 Century  
3 symmetric cells

大氣受熱力驅使，熱力受大氣控制  
Atmosphere is thermally forced, the heating mechanism are very complicated and motion dependent.

多重尺度擾動-平均流交互作用  
The dynamical response to the heating is also complicated because of eddies-mean flow interaction. The eddies are neither purely random nor purely regular.

25



$$\overline{v'T'} > 0$$

Dove 1837 weather is contributing; but was dismissed.

Bigelow 1902  
Defant 1921  
heat down-gradient transport

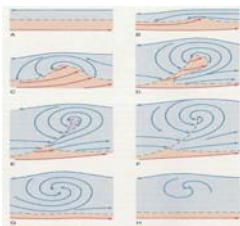
$$\overline{u'v'} > 0$$

Jefferey 1926 momentum up-gradient transport

Confirmed by Starr and J. Bjerkness 1948  
Starr: Physics of negative viscosity phenomena

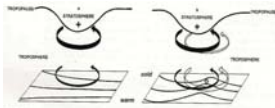
Origin of the eddies?? Baroclinic Instability  
Eady 1950, Charney 1959

26



Dove 1837 weather is contributing; but was dismissed.

V. Bjerkness 1937  
Symmetric component is unstable to the asymmetries.

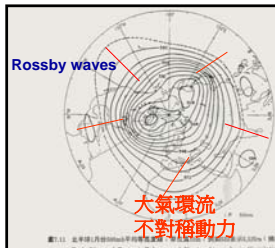


Baroclinic Instability  
Eady 1950, Charney 1959

高層低壓槽  
和低層低壓  
垂直交互作用

FIGURE 1.7 A schematic picture of development associated with the arrival of an upper-level positive PV anomaly, related to baroclinic dynamic instability, indicated by a westerly 'v' signet over a low-level baroclinic region. The circulation induced by the anomaly is indicated by the solid arrows, and potential temperature contours are shown at the lower boundary by thin lines. A low-level PV anomaly (the open 'v' signet in 8b), can also induce a cyclonic circulation, indicated by the open arrow in 8b, that acts to maintain the circulation pattern induced by the upper-level PV anomaly (Stokstad et al., 1985).

27



Rossby waves

Wave-mean flow interaction

$$\overline{v'T'} > 0$$

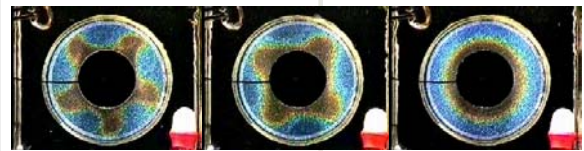
Heat down-gradient

$$\overline{u'v'} > 0$$

Momentum up-gradient

大氣環流  
不對稱動力

Baroclinic Instability (Stirring)



The Ultimate Problem in Meteorology Bjerknes 1911  
氣象的終極問題

I The Present state of the atmosphere must be characterized as accurately as possible. 正確的觀測大氣現狀  
[多重時空尺度]

II The intrinsic laws, according to which the subsequent states develop out of the preceding ones, must be known.  
正確的大氣運作規律

Numerical Weather Prediction 數值天氣預報  
[第一部電腦ENIAC, EBV model, 1950]  
The Observation component 觀測  
The diagnostic or analysis component 診斷分析  
The prognostic component 預報

Vilhelm Bjerknes (1862-1951)

科氏力 (18, 19)

Momentum Conservation (18)

$$\frac{\partial u}{\partial t} + u \frac{\partial u}{\partial x} + v \frac{\partial u}{\partial y} + w \frac{\partial u}{\partial z} - f v = -\frac{1}{\rho} \frac{\partial p}{\partial x} + \nu \nabla^2 u$$

$$\frac{\partial v}{\partial t} + u \frac{\partial v}{\partial x} + v \frac{\partial v}{\partial y} + w \frac{\partial v}{\partial z} + f u = -\frac{1}{\rho} \frac{\partial p}{\partial y} + \nu \nabla^2 v$$

Mass conservation (18)

$$\frac{\partial w}{\partial t} + u \frac{\partial w}{\partial x} + v \frac{\partial w}{\partial y} + w \frac{\partial w}{\partial z} = -\frac{1}{\rho} \frac{\partial p}{\partial z} - g + \nu \nabla^2 w$$

$$\frac{\partial \rho}{\partial t} + \frac{\partial u \rho}{\partial x} + \frac{\partial v \rho}{\partial y} + \frac{\partial w \rho}{\partial z} = 0$$

Energy conservation (19)

$$\frac{\partial \theta}{\partial t} + u \frac{\partial \theta}{\partial x} + v \frac{\partial \theta}{\partial y} + w \frac{\partial \theta}{\partial z} = Q$$

Equation of State (17, 18, 19)

$$p = \rho R_a T, \quad \theta = T \left( \frac{p_0}{p} \right)^{\frac{\gamma}{\gamma-1}}$$

Radiation 大氣輻射 (19, 20)  
Moisture Latent heat  
雲物理 (19, 20)

問蒼茫大氣，誰主浮沉？  
質量、動量、能量與大氣狀態方程式

30

## Lewis Fry Richardson, 1881–1953.



L. F. Richardson, 1921

During WWI, Richardson computed by hand the pressure change at a single point.

It took him **two years** !

His 'forecast' was a catastrophic failure:

$$\Delta p = 145 \text{ hPa in 6 hours}$$

His **method** was unimpeachable.

So, *what went wrong?*

Peter Lynch

RICHARDSON GRID

$$\frac{df}{dx} \rightarrow \frac{f(x + \Delta x) - f(x - \Delta x)}{2\Delta x}$$

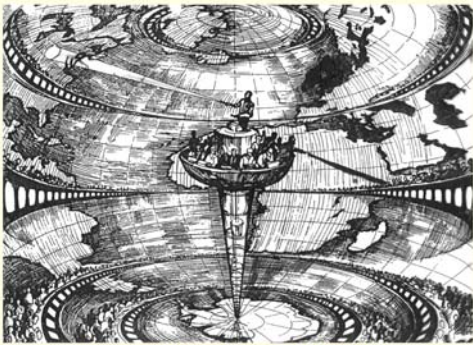
$$\frac{dQ}{dt} \rightarrow \frac{Q^{n+1} - Q^{n-1}}{2\Delta t} = F^n$$

**13x13=169個ODE**

**169 自由度**

32

## Richardson's Dream



Richardson's Forecast Factory (A. Lammert), Dagbladet Nyheter, Stockholm. Reproduced from L. Bengtsson, *ECMWF*, 1984

**64,000 Computers: The first Massively Parallel Processor**

## first weather forecast - ENIAC, 1950



In front of the Eniac, Aberdeen Proving Ground, April 4, 1950, on the occasion of the first numerical weather computations carried out with the aid of a high-speed computer. <sup>34</sup>

## The ENIAC Electronic Numerical Integrator and Computer



18000 vacuum tubes  
70000 resistors  
10000 capacitor  
6000 switches

140 K Watts power

No high-level language  
Assembly language

**500 Flops**  
**Function Table 0.001 s**

**3,700,000,000** times slower than current day large computer

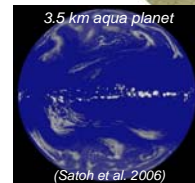
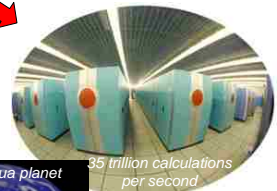
第一部電腦 氣象預報

35

## ENIAC - late 40s



## Earth Simulator -- 2002



NASDA, JAERI, JAMSTEC

(Satoh et al. 2006)

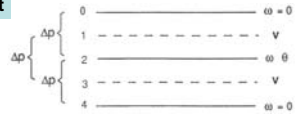
36

mid 50

Observational study of eddy transport **UCLA, MIT**  
 Laboratory study **U of Chicago, MIT**  
 NWP — Charney, Fjortoft and Neumann (1950) **E Barotropic**  
 Charney and Phillips (1953) **2 level QG model**

**Phillips' numerical experiment**

**Phillips (1956)**  
**two-level**  
**Quasi-Geostrophic model**



Corresponding to (IV.3.1) for  $\ell = 1, 2$  and (IV.3.2) for  $\ell = 1$  but with a new vertical index as shown in Fig. IV.8, the equations of his model are

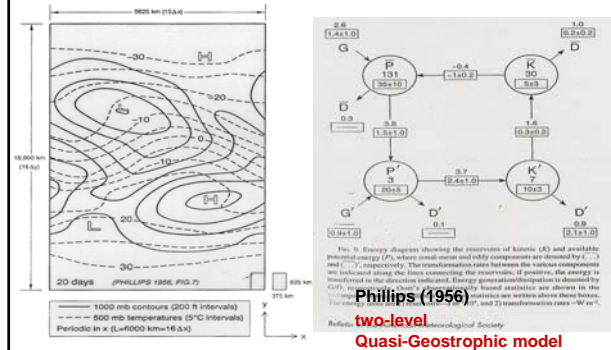
$$\left(\frac{\partial}{\partial t} + v_{g1} \cdot \nabla\right) (\zeta_{g1} + f) - \zeta_{g1} \frac{\omega_2}{\Delta p} = A \nabla^2 \zeta_{g1}, \quad (IV.4.1)$$

$$\left(\frac{\partial}{\partial t} + v_{g3} \cdot \nabla\right) (\zeta_{g3} + f) + \zeta_{g3} \frac{\omega_2}{\Delta p} = A \nabla^2 \zeta_{g3} - k \zeta_{g3}, \quad (IV.4.2)$$

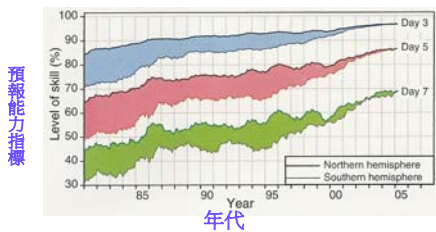
$$\left(\frac{\partial}{\partial t} + v_{g1} \cdot \nabla\right) (\psi_1 - \psi_3) - \frac{1}{\mu^2} \frac{\partial \omega_2}{\partial p} = A \nabla^2 (\psi_1 - \psi_3) + \frac{R}{c_p} \frac{1}{\zeta_{g3}}, \quad (IV.4.3)$$

**to study meteorology as an experimental science**

**To study meteorology as an experimental science**



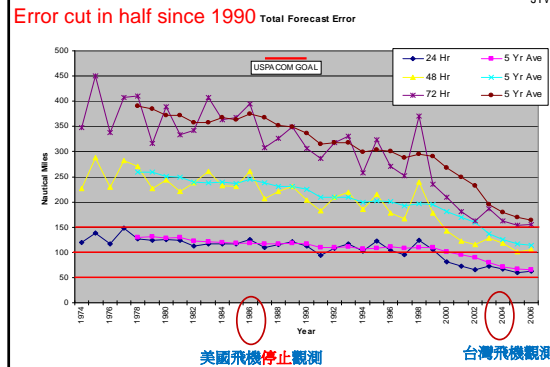
南北半球 對於3, 5, 7天之**預報能力**隨時間的進展



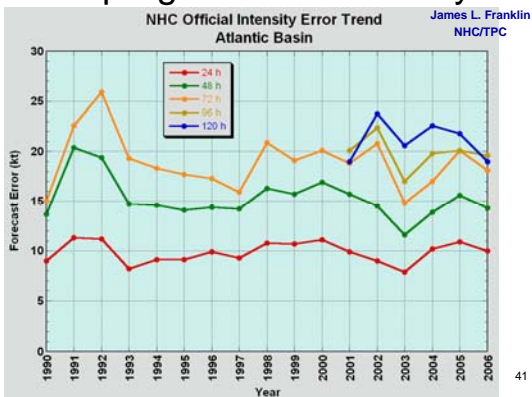
南北差異日漸減少主要是由於近年來**衛星觀測**以及**資料同化**技術日漸成熟  
**7天預報一年進步約1.5%，3天預報一年進步0.3%**

**West Pac Track Errors**

Edward Fukada  
 JTWC

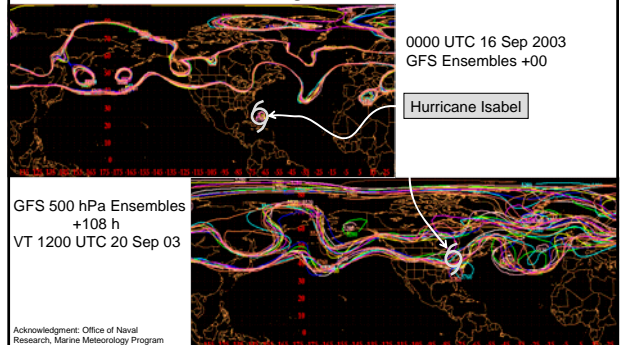


**No progress with intensity**



**The Downstream Influences of the Extratropical Transition of Tropical Cyclones**

Patrick Harr  
 Naval Postgraduate School



Acknowledgment: Office of Naval Research, Marine Meteorology Program

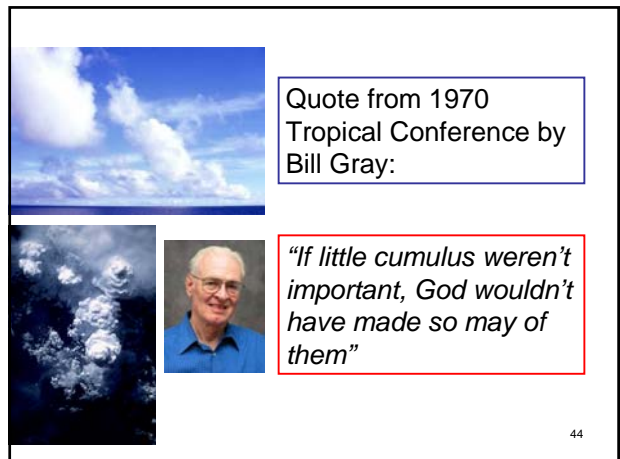
### 颱風潛熱與其它能量的比較

賀伯颱風的全台灣平均總雨量為400mm  
 $400\text{ mm} = 0.4\text{ m}$   
 $0.4\text{ m} * 1000\text{ kg m}^{-3} * 2.5 * 10^6\text{ J kg}^{-1}$   
 $= 10^9\text{ J m}^2$   
 $10^9\text{ J m}^2 * 3.5 * 10^{10}\text{ m}^2$   
 $= 3.5 * 10^{19}\text{ J} \sim 10^{20}\text{ J}$

${}^1_0\text{n} + {}^{235}_{92}\text{U} \rightarrow {}^{142}_{56}\text{Ba} + {}^{91}_{36}\text{Kr} + 3 {}^1_0\text{n}$

$1.68 * 10^{13}\text{ J/mol}$   
 $\Rightarrow 1.46 * 10^6\text{ kg U}^{235} (6 * 10^6\text{ mol})$

能量估計值		備註
賀伯颱風降雨總潛熱能量	$10^{20}\text{ J}$	可使台灣整層大氣增溫100度
台灣一年用電量	$5 * 10^{17}\text{ J}$	需數百年用電量才相當
全世界核子彈爆炸釋放能量	$2 * 10^{19}$ $\sim 2 * 10^{20}\text{ J}$	與賀伯颱風同等級
核戰後燃燒釋放能量	$2 * 10^{20}\text{ J}$	與賀伯颱風同等級
地球一天接受的太陽能量	$1.5 * 10^{22}\text{ J}$	數百個賀伯颱風
Tunguska隕石撞地球 (西元1908年, 西伯利亞)	$10^{16}\text{ J}$	賀伯颱風的萬分之一
火流星撞地球 (恐龍滅絕?)	$4 * 10^{23}\text{ J}$	數千個賀伯颱風



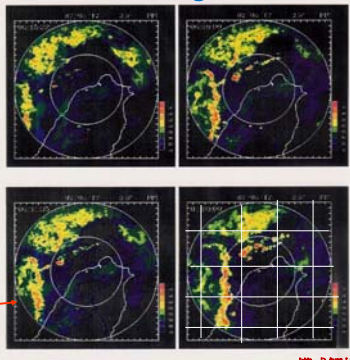
Quote from 1970 Tropical Conference by Bill Gray:

*"If little cumulus weren't important, God wouldn't have made so many of them"*

### Mesoscale Organization

狂風不終朝  
 暴雨不終日  
 孰為此者  
 天地  
 天地尚不能久  
 而況人乎

飆線  
 Squall line



1987年5月17日鋒前飆線之雷達回波圖 模式解析?

### Numerical Modeling of a Squall Line

(Houze 1977)

Mesoscale Organization

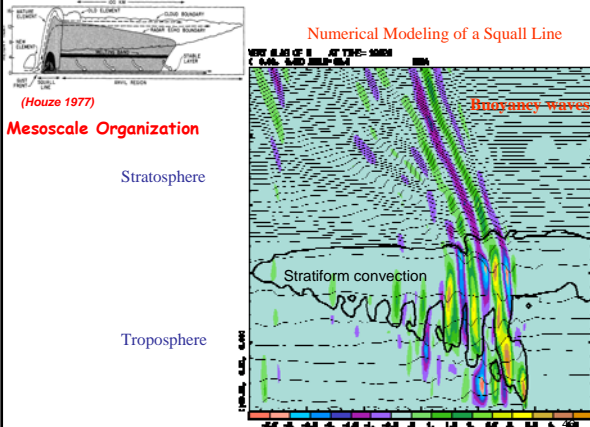
Stratosphere

Troposphere

Waves

Instability

Gravity waves

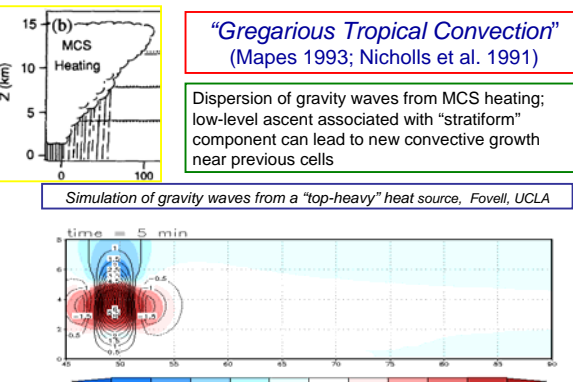


### Coupling Between Convection and Its Environment

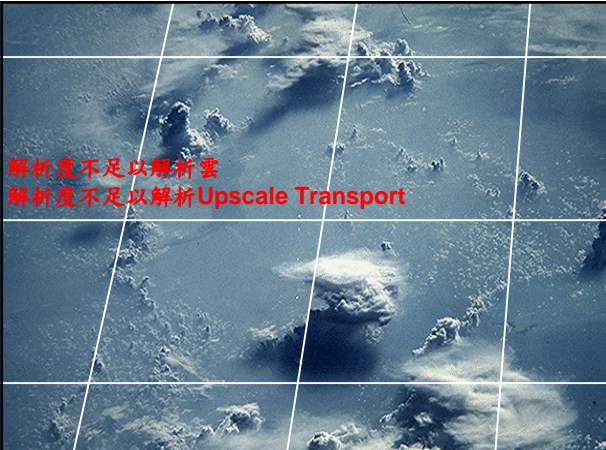
"Gregarious Tropical Convection" (Mapes 1993; Nicholls et al. 1991)

Dispersion of gravity waves from MCS heating; low-level ascent associated with "stratiform" component can lead to new convective growth near previous cells

Simulation of gravity waves from a "top-heavy" heat source, Fovell, UCLA



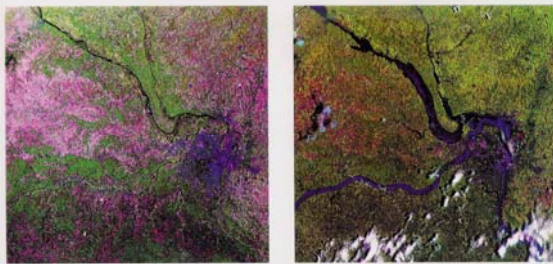
解析度不足以解析雲  
 解析度不足以解析Upscale Transport





風雨之不時，是無世而不常有之。 荀子天論

1988 年際變化 季節預報 1993



Heavy rains in the summer of 1993 produced floods along most of the Mississippi River in the central United States, as shown in these Earth satellite photographs of St. Louis, Missouri on July 4, 1988 (left) and July 13, 1993 (right). Extreme climatic events may be increasing in frequency as a consequence of added radiative absorbing gases in the atmosphere.

49

The Atmosphere is Moist 濕的大氣

Water vapor is an efficient absorber and emitter of Long-wave radiation. [Green House Effect.]

溫室效應 大氣輻射

Water vapor stores energy in the form of "latent heat" [Evaporative cooling of surface.]

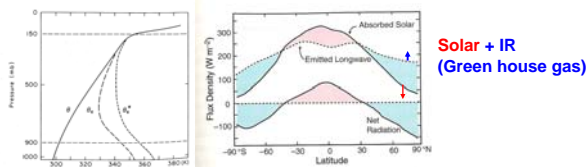
地球表面蒸發冷卻

Water vapor can condense release latent heat. [A variety of cloud and associated processes.]

雲過程

50

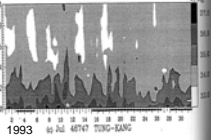
Conditional unstable atm.



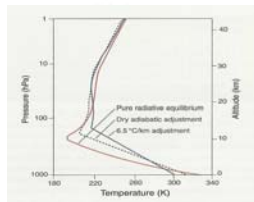
Solar + IR (Green house gas)

Finite amplitude forcing Rooted in BL High h\* for instability

Deep Convection



東港 7月



Radiative Convective Adjustment 1K/day cooling 1m/year precipitation 水文循環

大氣浮力 阿基米得原理

濕絕熱降溫率 Moist adiabatic lapse rate

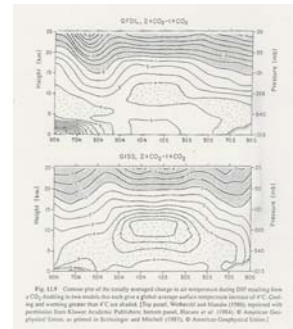
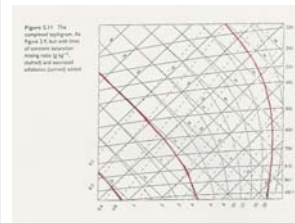


Fig. 12.4 Change plot of the locally averaged change in air temperature during 2000 matching years of CO2 doubling to two models that have given a global average and temperature decrease of 4°C. Cooling and warming greater than 4°C are shaded. (This panel, National Oak Institute (1986), reprinted with permission from Elsevier Academic Publishers, Amsterdam, Boston, MA, 1986. © American Geophysical Union, as printed in Stouffer and Mitchell (1987), © American Geophysical Union.)

52

Harteman, Climate text book

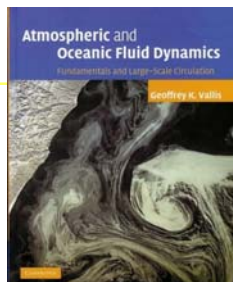
20th Century

Geophysical Fluid Dynamics (GFD) Atmospheric Oceanic Fluid Dynamics (AOFD)

is for those interested in doing research in the physics, chemistry, and/or biology of Earth fluid environment.



Fig. 9.2 Karman vortex streets in (a) the laboratory, for water flowing past a cylinder [From M. Van Dyke, An Album of Fluid Motion, Parabolic Press, Stanford, Calif. (1982) p. 56.], and (b) in the atmosphere, for a cumulus-topped boundary layer flowing past an island [NASA MODIS imagery].



熱力學 + 流體力學

Euler 1755 
$$\frac{d}{dt} \int_{v_m} \rho \vec{v} \, dv = - \int_{\partial v_m} p \, d\vec{s}$$

$$\int_{v_m} \rho \frac{d\vec{v}}{dt} \, dv = - \int_{v_m} \nabla p \, dv$$

$$\rho \frac{d\vec{v}}{dt} = - \nabla p$$

Lagrange 1781

$$\frac{\partial \vec{u}}{\partial t} + \underbrace{\vec{\zeta} \times \vec{u}}_{\text{Rotation, Vortex}} = - \frac{1}{\rho} \nabla p - \nabla K - \nabla \Phi$$

Rotation, Vortex

Lorentz Force Law

$$\mathbf{F} = q(\mathbf{E} + \mathbf{v} \times \mathbf{B})$$

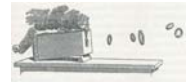
$$\mathbf{F} = q(-\nabla V + \mathbf{v} \times \mathbf{B})$$

Helmholtz 1858

$$\frac{\partial \vec{\zeta}}{\partial t} + \vec{v} \cdot \nabla \vec{\zeta} + \vec{\zeta} \cdot \nabla \vec{v} = \vec{\zeta} \cdot \nabla \vec{v} + \vec{B}$$

$$\vec{B} = \nabla \times \left( - \frac{1}{\rho} \nabla p \right)$$

19世紀煙圈表演



**Wake Turbulence**

Fig. 8.9. Vortices trailing from the wingtips of a Boeing 727. Figure courtesy of NASA.

Fig. 8.10. Sketch of the flow along an airfoil. The wing is shown in grey, the contour  $C'$  is shown by the thick solid line.

Fig. 8.11. Two boats carrying sails with very different aspect ratios.

**Biomath**

Figures 3.102 (a) A horseshoe vortex representing a wing with a uniform lift distribution. (b) Lift distribution on an elliptic wing.

Figure 3.104 The vortices in the wake of an oscillating wing. Idealized under the assumption that the lift fluctuation is very small so that the direction of the wake due to the vortices in it is also very small.  $L$  is the lift,  $\alpha$  is the angle of attack.

Figure 3.105 The vortex wake behind a stork in level flight.

Y. C. Fung 56

Guinn and Schubert 1994

Hendricks et al. (2004)

**Vortical Hot Towers**

- 10 km in diameter on average

1969/08/16 Debbie

A : increase its diameters from 50 to 100 km in ~2 hours.

C, D : grew into circular clouds with about a 20 km diameter in less than 15 min.

$$\frac{DP}{Dt} = \frac{\partial P}{\partial t} + u \frac{\partial P}{\partial x} + v \frac{\partial P}{\partial y} + w \frac{\partial P}{\partial z} \quad P = \frac{\bar{\zeta} \cdot \nabla \theta}{\rho} \quad \text{Potential Vorticity}$$

$$\frac{D\zeta}{Dt} = \frac{\partial \zeta}{\partial t} + u \frac{\partial \zeta}{\partial x} + v \frac{\partial \zeta}{\partial y} \quad \text{Barotropic Vorticity}$$

$$\frac{D\zeta_g}{Dt} = \frac{\partial \zeta_g}{\partial t} + u_g \frac{\partial \zeta_g}{\partial x} + v_g \frac{\partial \zeta_g}{\partial y} \quad \text{Quasi-Geostrophic Vorticity (Strong Rotation)}$$

$$\zeta_g = \nabla^2 \psi \quad u_g = -\frac{\partial \psi}{\partial y} \quad v_g = \frac{\partial \psi}{\partial x} \quad \nabla \cdot \vec{V} = 0 \quad \text{Nondivergent Barotropic Equation}$$

$$\frac{\partial \zeta}{\partial t} + \frac{\partial(\psi, \zeta)}{\partial(x, y)} = \nu \nabla^2 \zeta$$

59

**2D Turbulence**

**Stratification and/or Rotation**


**Vortex Waves Turbulence**

$$\frac{\partial \zeta}{\partial t} + u \frac{\partial \zeta}{\partial x} + v \frac{\partial \zeta}{\partial y} = \nu \nabla^2 \zeta$$

$$u = -\frac{\partial \psi}{\partial y}, \quad v = \frac{\partial \psi}{\partial x}$$


$$\frac{\partial \zeta}{\partial t} + \frac{\partial(\psi, \zeta)}{\partial(x, y)} = \nu \nabla^2 \zeta$$

60

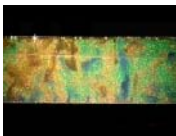


旋轉  
Rotation

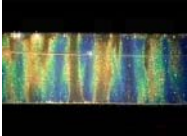
Coriolis Force  
Non-inertial Frame



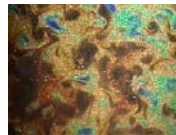
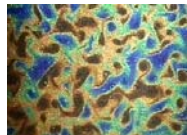
3D



2D (strong rotation)



Taylor columns Vortex Tubes



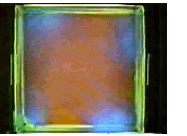



Vortices with sharp edge

62  
Kyoto Univ. GFD group

Waves with zero potential vorticity

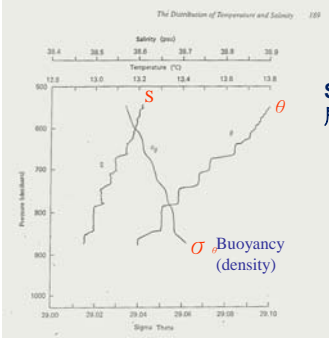
Non-rotation      rotation      rotation

Gravity waves      Kelvin Waves  
Edge waves

63

Ocean Spice



Stratification  
層化

$\sigma$  Buoyancy (density)

64


Kelvin-Helmholtz Instability  
Turbulent Mixing 晴空亂流 Intermediate wavelength






65

Multiple Scale Interactions in Vortex



Wave mean flow interaction in stable stratified fluid  
Turbulent feed back to the vortex mean flow

2D turbulence

66

$$\frac{d}{dt} \int E(k) dk = 0, \quad \frac{d}{dt} \left( \int k^2 E(k) dk \right) = \frac{d}{dt} \int Z(k) dk = 0$$

$$\frac{d}{dt} \left( \int (k - k_1)^2 E(k) dk \right) > 0$$

$$\frac{d}{dt} \left( \int k^2 E(k) dk + k_1^2 \int E(k) dk - 2k_1 \int k E(k) dk \right) > 0$$

$$\frac{d}{dt} \left( \frac{\int k E(k) dk}{\int E(k) dk} \right) < 0, \quad \text{Kinetic energy moves toward large scales}$$

$$\frac{d}{dt} \left( \int (k^2 - k_1^2)^2 E(k) dk \right) > 0$$

$$\frac{d}{dt} \left( \int k^2 Z(k) dk + k_1^4 \int E(k) dk - 2k_1^2 \int k^2 E(k) dk \right) > 0$$

$$\frac{d}{dt} \left( \frac{\int k^2 Z(k) dk}{\int Z(k) dk} \right) > 0, \quad \text{Enstrophy moves toward small scales}$$

**Non-divergent barotropic model (Nearly Inviscid Fluid)**

$$\frac{\partial}{\partial t} \zeta + J(\psi, \zeta) = \nu \nabla^2 \zeta \quad \nabla^2 \psi = \zeta$$

The energy and enstrophy relations


$$\frac{dE}{dt} = -2\nu Z, \quad E = \iint \frac{1}{2} (u^2 + v^2) dx dy \quad \text{kinetic energy}$$

$$\frac{dZ}{dt} = -2\nu P, \quad Z = \iint \frac{1}{2} \zeta^2 dx dy \quad \text{enstrophy}$$


$$P = \iint \frac{1}{2} \nabla \zeta \cdot \nabla \zeta dx dy \quad \text{palinstrophy}$$

Batchelor 1969 68

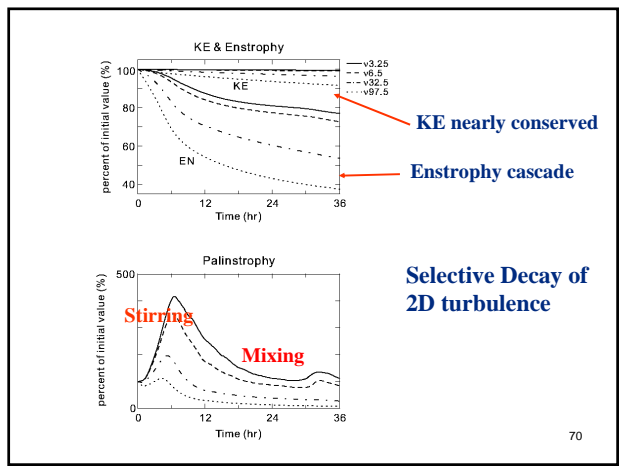
Small viscosity led to large palinstrophy and the large enstrophy cascade



**Stirring**



69



$E \sim p^2 / L^2$  (KE)    geostrophy

$Z \sim p^2 / L^4$  (Enstrophy)

KE nearly conserved     $L \sim p^{-1}$

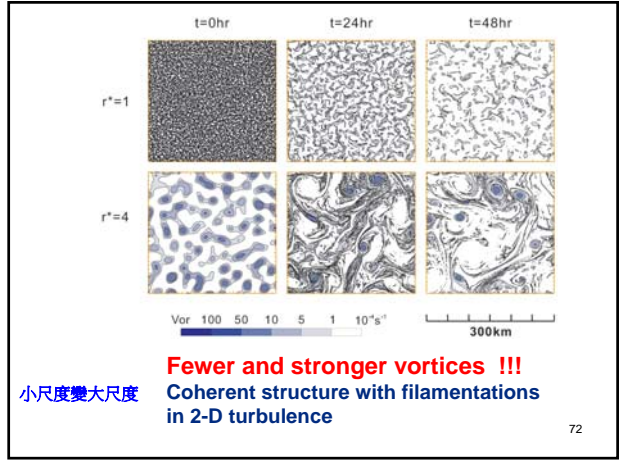
Enstrophy cascade     $L \uparrow$  (L increase    Z decrease)

Selective Decay of 2D turbulence

The vortices become, on the average, larger, stronger, and fewer.

Merger and Axisymmetrization Dynamics

71



Weiss(1981,1991), Rozoff et al. (2004)

$$\frac{D}{Dt}(\nabla\zeta) = -J(\nabla\psi, \zeta)$$

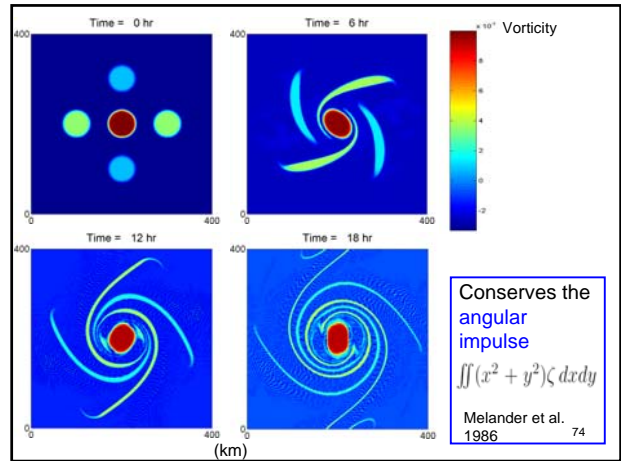
$$\rightarrow \nabla\zeta(t) \propto \exp(\lambda t) \quad \lambda = \pm \frac{1}{2}\sqrt{Q} = \pm \frac{1}{2}\sqrt{S_1^2 + S_2^2 - \zeta^2}$$

$$S_1 = \frac{\partial u}{\partial x} - \frac{\partial v}{\partial y} \quad (\text{stretch deformation})$$

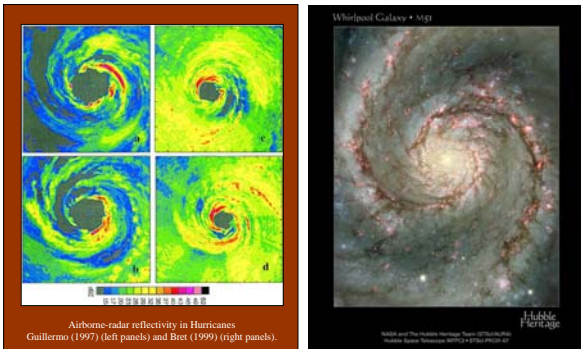
$$S_2 = \frac{\partial v}{\partial x} + \frac{\partial u}{\partial y} \quad (\text{shear deformation})$$

- Q > 0 (strain dominates)
  - vorticity gradient will be stretched
- Q < 0 (vorticity dominates)
  - vortex is stable (survival of eyewall meso-vortices)

73



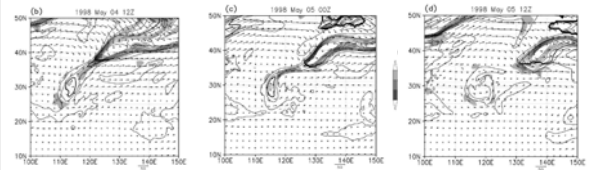
Spiral Band in Hurricane and Galaxy



Kossin and Schubert 2001

75

Filamentation Time Diagnosis of Thinning Troughs and Cutoff Lows



$$\tau_{fil} = 2 \left\{ \delta + \left[ S_n^2 + S_s^2 - \left( \zeta - \frac{2u \tan \phi}{a} \right)^2 \right]^{1/2} \right\}^{-1} \quad \text{for } S_n^2 + S_s^2 > \left( \zeta - \frac{2u \tan \phi}{a} \right)^2$$

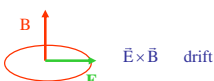
Tsai et al. 2009

Electron density redistribution in experimental plasma physics

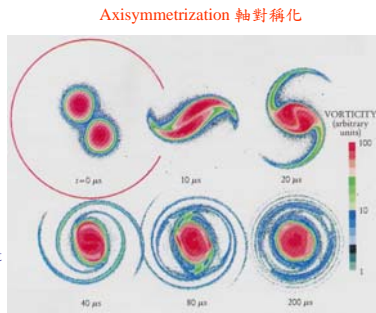
single sign charge  
+  
axial magnetic field  
confinement

$$E = -\nabla\psi$$

$$\nabla \cdot E = -\nabla^2\psi = \frac{\rho}{\epsilon}$$



Coriolis force



Core is protected, thin filaments from edges

77

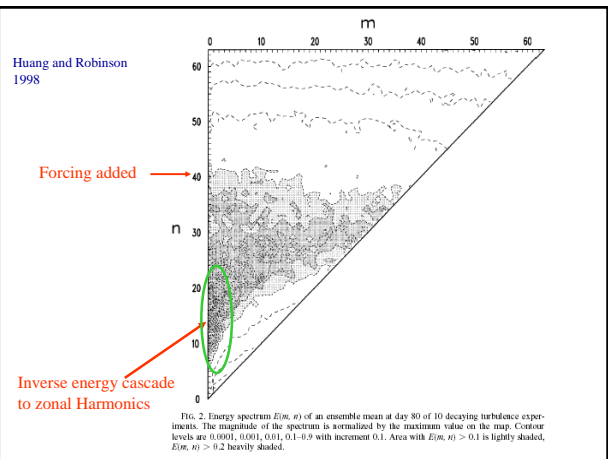
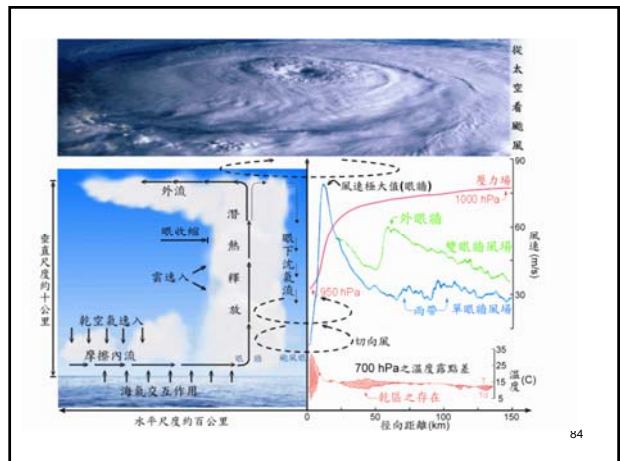
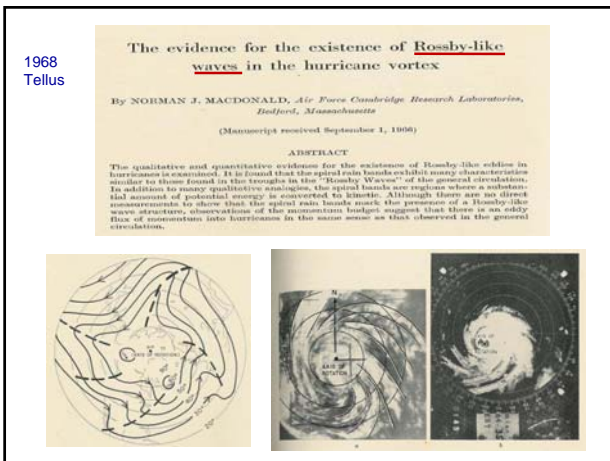
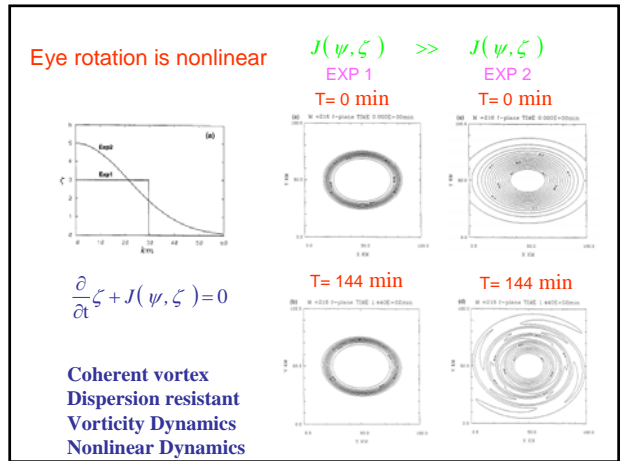
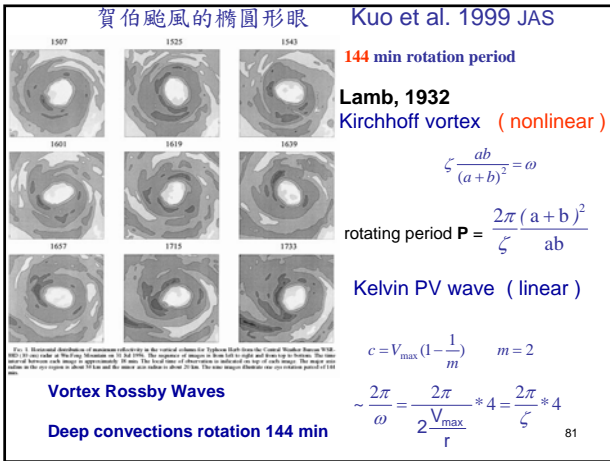
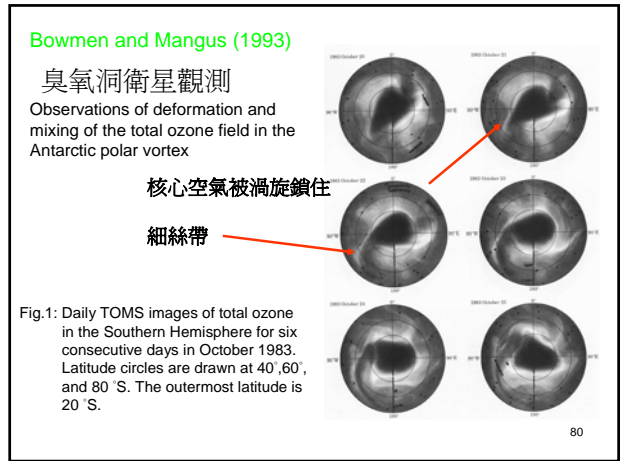
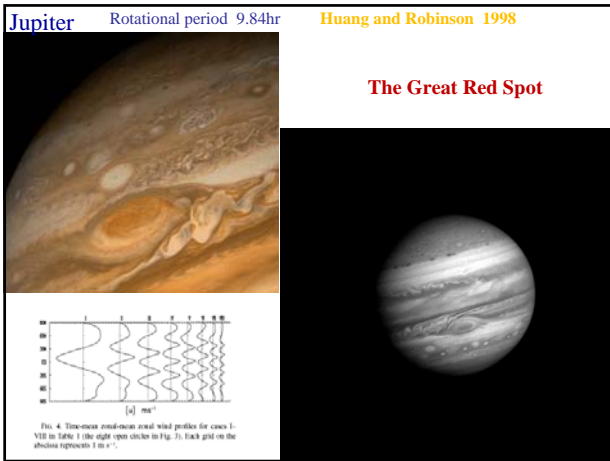


FIG. 2. Energy spectrum  $E(m, n)$  of an ensemble mean at day 30 of 10 decaying turbulence experiments. The magnitude of the spectrum is normalized by the maximum value on the map. Contour levels are 0.0001, 0.001, 0.01, 0.1-0.9 with increment 0.1. Area with  $E(m, n) > 0.1$  is lightly shaded,  $E(m, n) > 0.2$  heavily shaded.

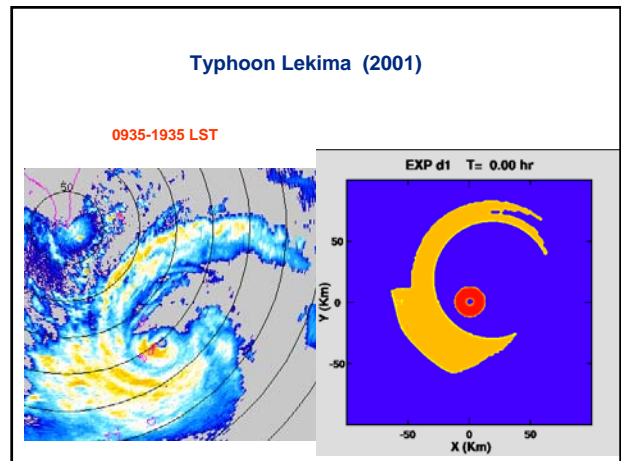
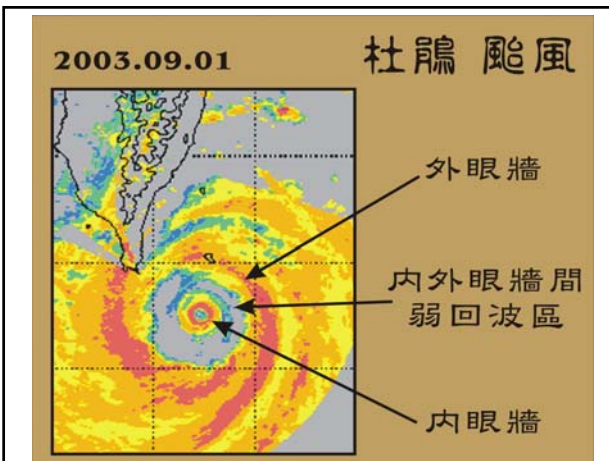
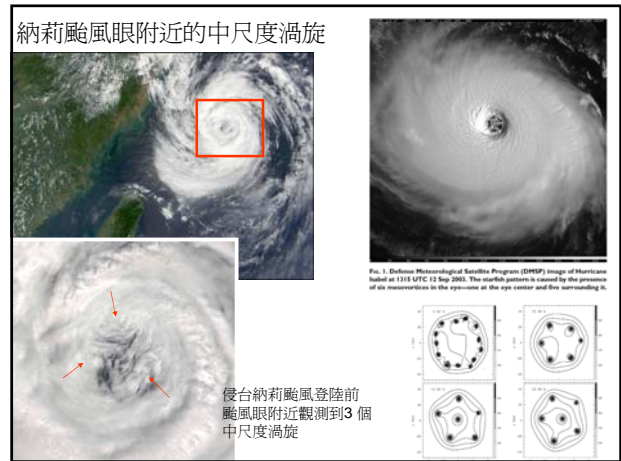
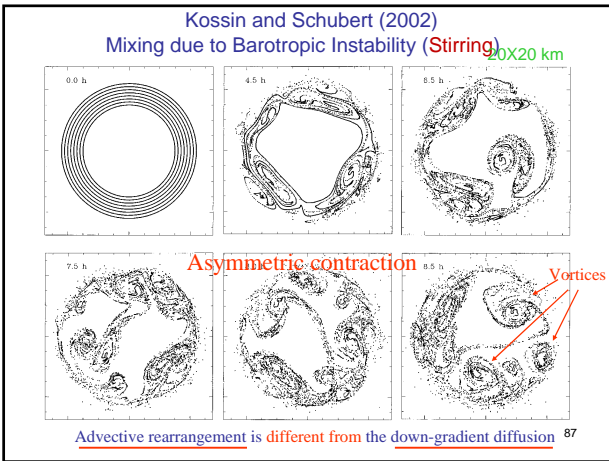
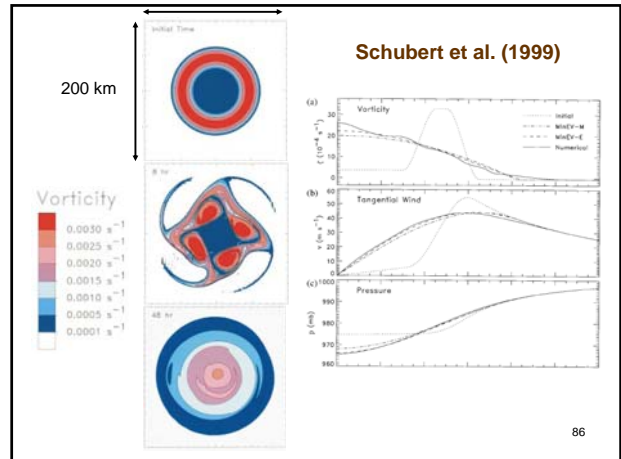
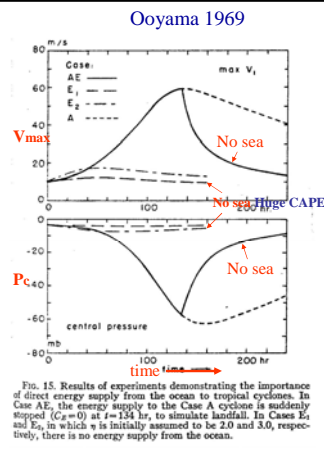


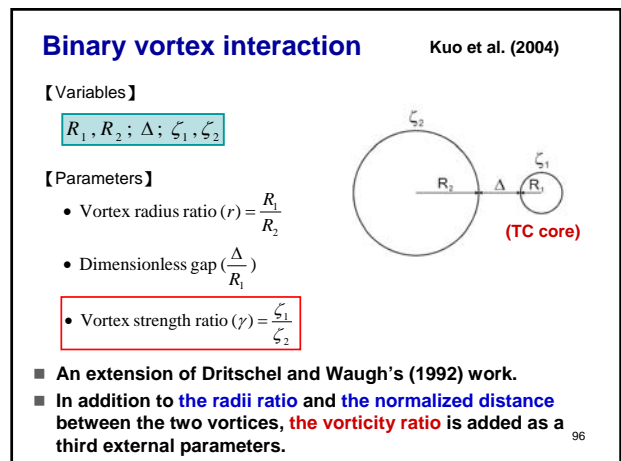
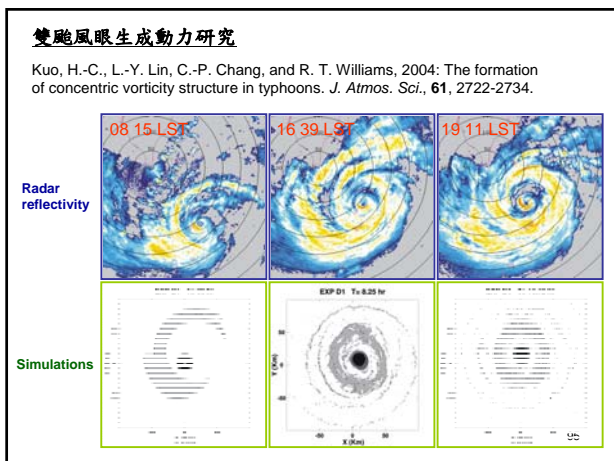
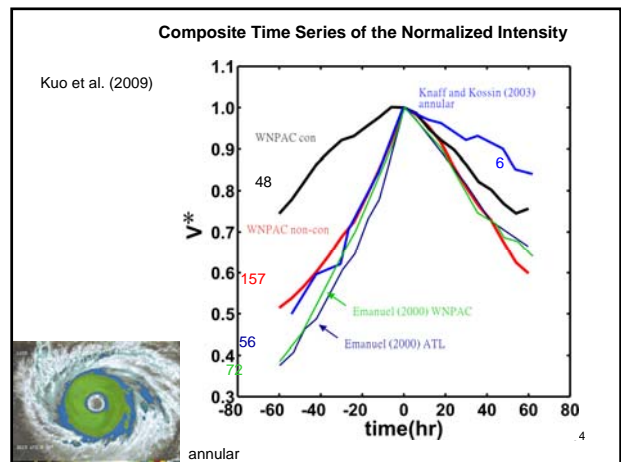
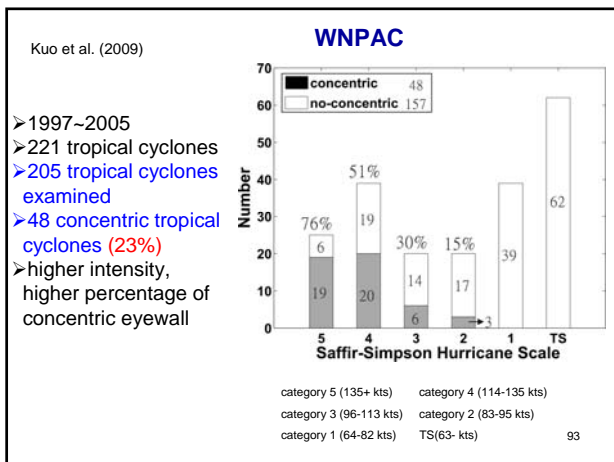
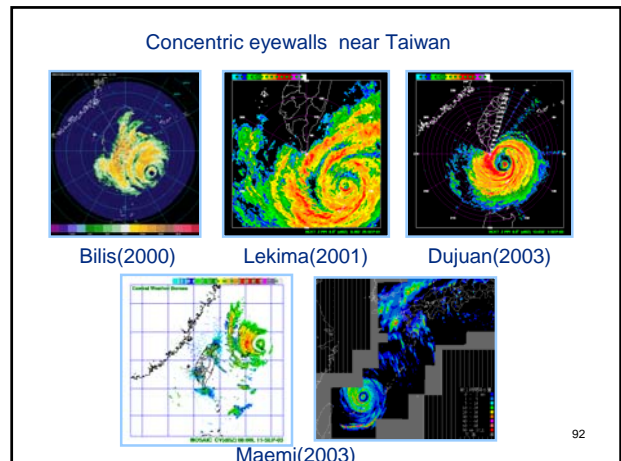
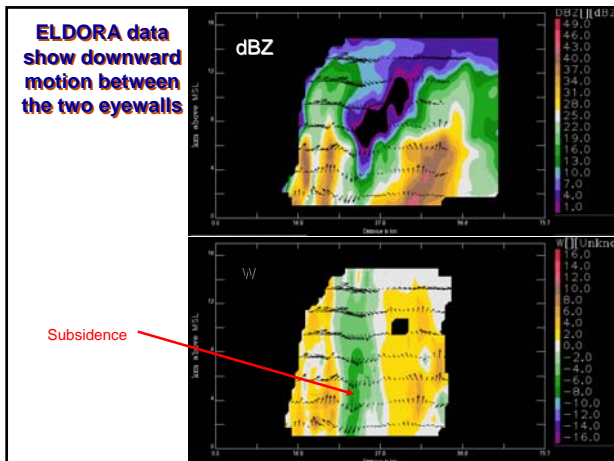
The important of air-sea interaction to the maintenance of tropical cyclone.

“Tritium measurement indicates 60%-80% of the water in the eyewall could have evaporated from the ocean not so long ago.”

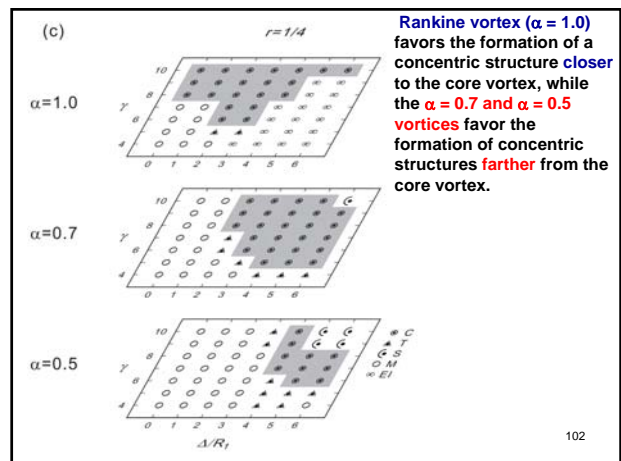
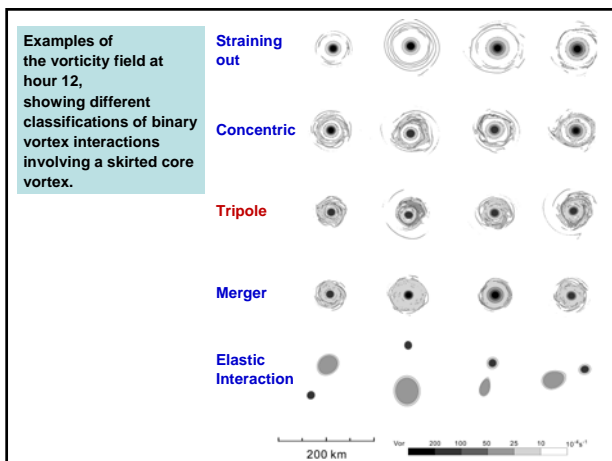
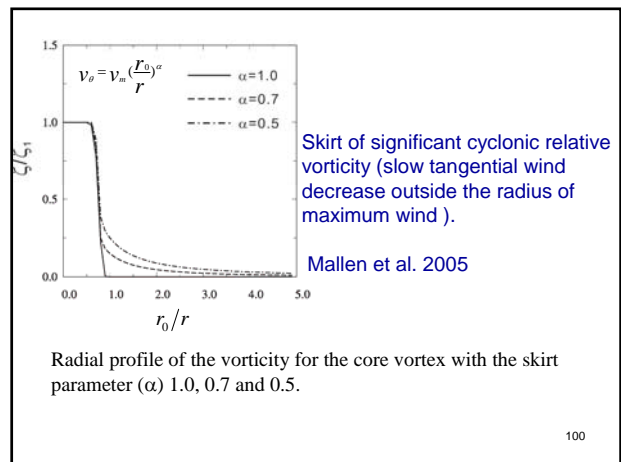
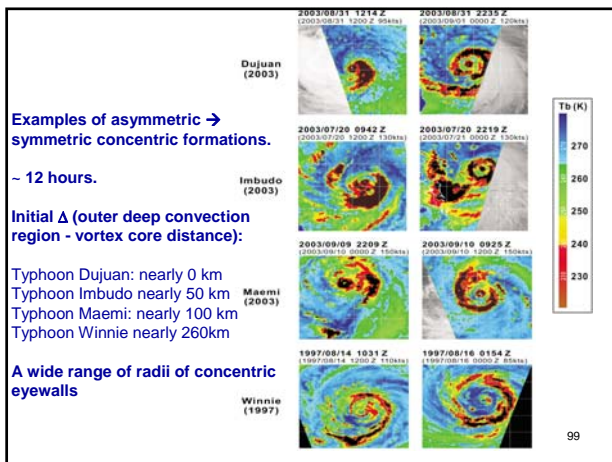
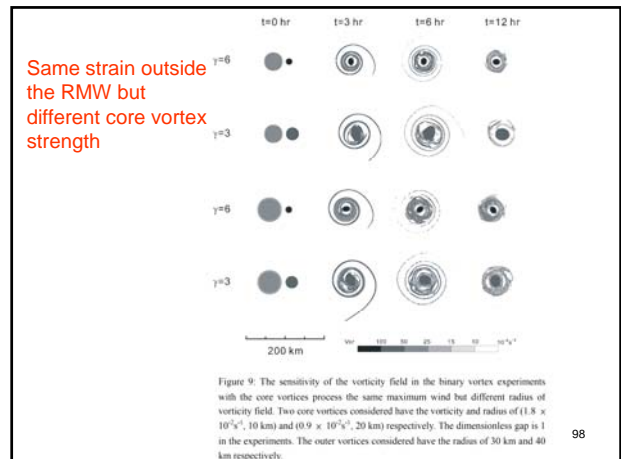
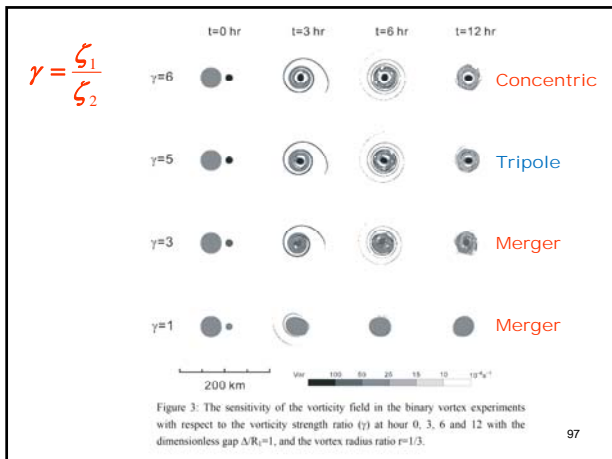
“Tropical cyclone consumes a great deal more latent heat than it can collect from the pre-existing atmospheric vapor.”

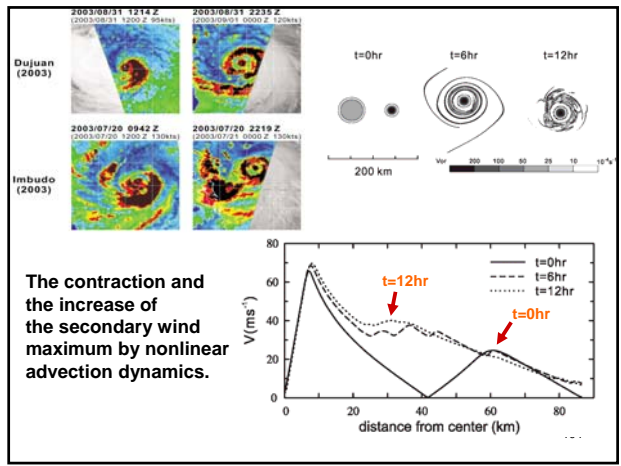
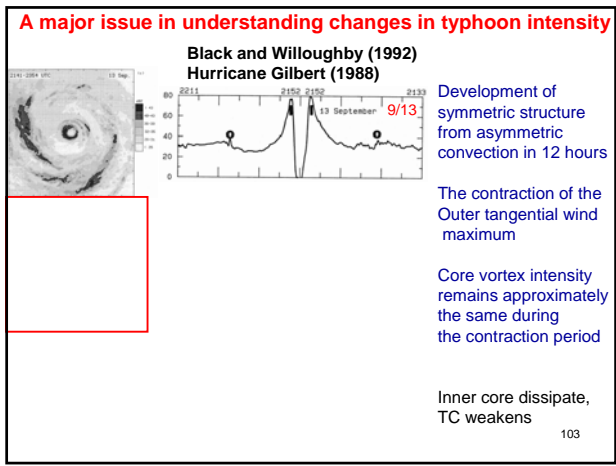
Huge CAPE without air-sea interaction make NO tropical cyclone.











Terwey and Montgomery, June JGR 2008

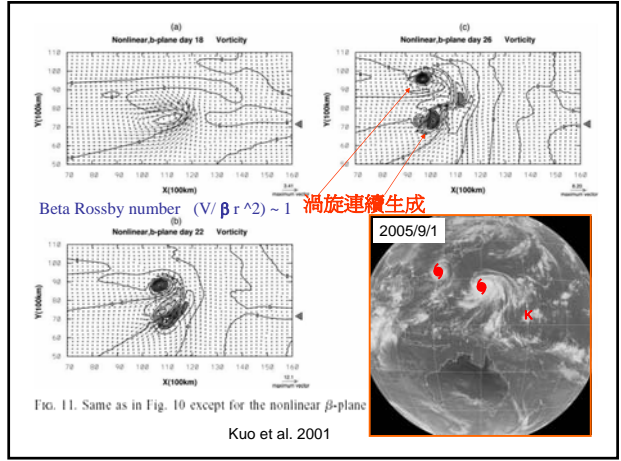
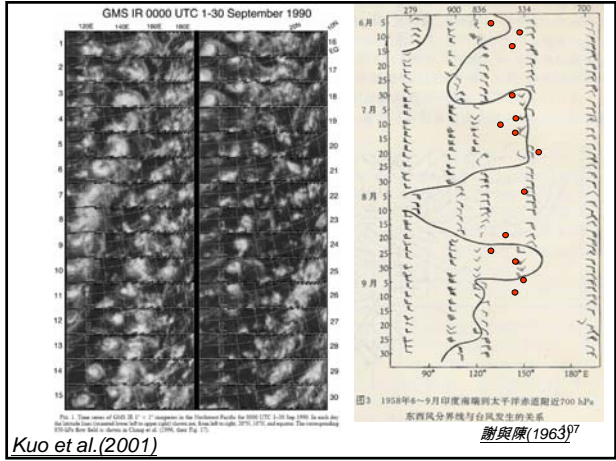
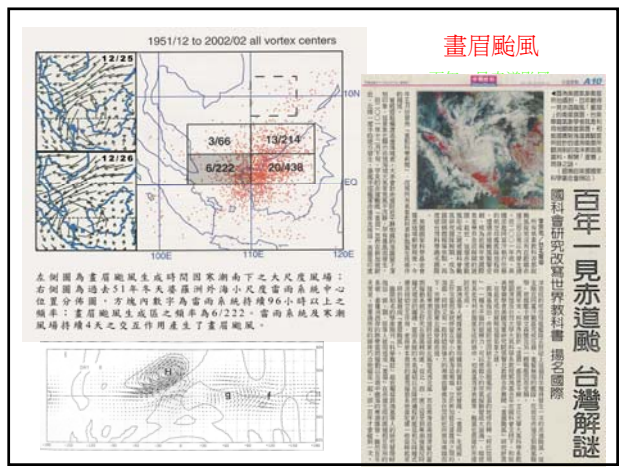
D12112 TERWEY AND MONTGOMERY: MODELED SECONDARY EYEWALL FORMATION D12112

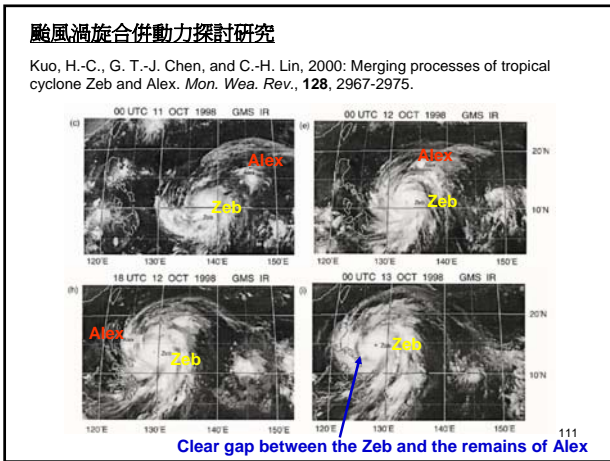
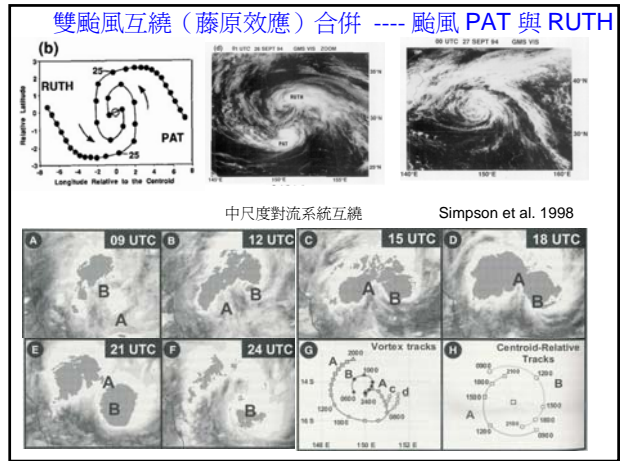
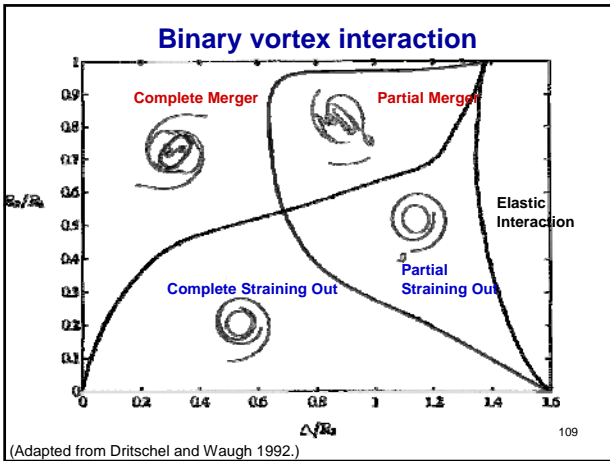
Table 1. List of Secondary Eyewall Formation Hypotheses With Summary of Relevance to Our Modeled Hurricanes\*

Authors	Hypothesis Summary	Relevance to Current Model Results	Type
Willoughby et al. [1982] line research of Zipser [1977] Willoughby [1979]	Downdrafts from the primary eyewall force a ring of convective updrafts. Internal resonance between local inertia period and asymmetric friction due to storm rotation.	Few downdraft-forced updrafts during this time in the simulations. No systematic storm motion in the simulated storms.	O
Hoskins [1983]	Topographic effects	No topographic forcing in the simulations.	O
Willoughby et al. [1984]	Ice microphysics	"Warm-rain" (no-ice) sensitivity case also produces secondary eyewall.	A
Melinaro and Shabli [1983] and Melinaro and Hildebrand [1989]	Synoptic-scale forcings (e.g., inflow surges, upper-level momentum fluxes)	No synoptic-scale forcings in the simulations.	O
Montgomery and Kallenbach [1997], Comp and Montgomery [2001] and Terwey and Montgomery [2003]	Internal dynamics-assignment via tilted vortex Rossby wave processes, collection of wave energy near stagnation or <b>cut-off</b>	Possible explanation	N
<b>Nong and Lau [2003]</b>	Sustained eddy momentum fluxes and WISHE feedback	Possible explanation	A
<b>Kuo et al. [2004, 2008]</b>	Asymmetrization of positive vorticity perturbations around a strong and tight core of convection	Possible explanation	N

\*The type column refers to the type of model or observations that were used to formulate the hypothesis. O stands for observationally-based, A stands for asymmetric model, N stands for nonasymmetric model.

1Ub





### Multiple Scale Interactions!

- **Waves, Turbulence, and Vortices**  
波動、亂流、渦旋 (2D turbulence)
- **Cumulus Parameterization**  
積雲合成作用 (3D turbulence)

Concerning turbulence, Sir Horace Lamb is quoted in an address to the British Association for the Advancement of Science as follow:  
*I am an old man now, and when I die and go to Heaven there are two matters on which I hope for enlightenment. One is quantum electrodynamics, and the other is the turbulent motion of fluids. About the former I am rather optimistic.*

**Before too long, climate modelers will have much more in common with turbulence modelers and micrometeorologists than either group now seems to realize.**

Tennekes 1978, *Turbulent Flow in Two and Three Dimension*  
 (Tennekes and Lumley, *A first course in turbulence*)

### Spectral Method $f = \sum a_n \phi_n$

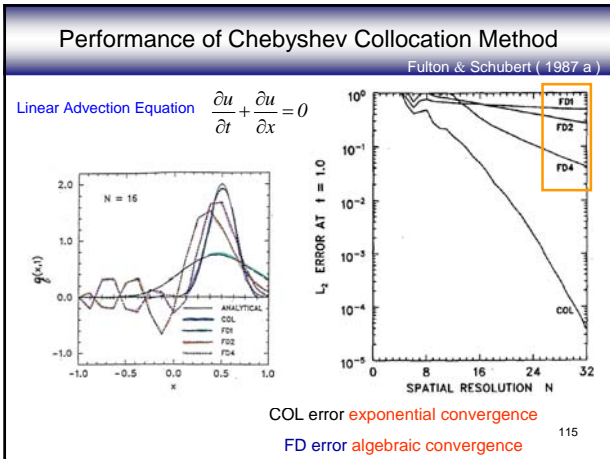
1. Completeness (完整性) → Sufficient condition
2. Orthogonality (正交性)
3. Speed of convergence (收斂速度)
4. Fast Transform (快速轉換)

Application

#### Sturm-Liouville equation

$$L\phi(x) = -\frac{d}{dx}(p(x)\phi'(x)) + q(x)\phi(x) = \lambda W(x)\phi(x)$$

with suitable boundary conditions and restrictions on functions  $p(x)$ ,  $q(x)$ ,  $W(x)$ , we have a countably infinite set of solutions  $\{\phi_n(x)\}_{n=0}^{\infty}$  corresponding to discrete eigenvalues  $\{\lambda_n\}_{n=0}^{\infty}$



### Shallow Water Equation

$$\frac{\partial u}{\partial t} + u \frac{\partial u}{\partial x} + v \frac{\partial u}{\partial y} - fv + \frac{\partial h}{\partial x} = 0 \quad (1)$$

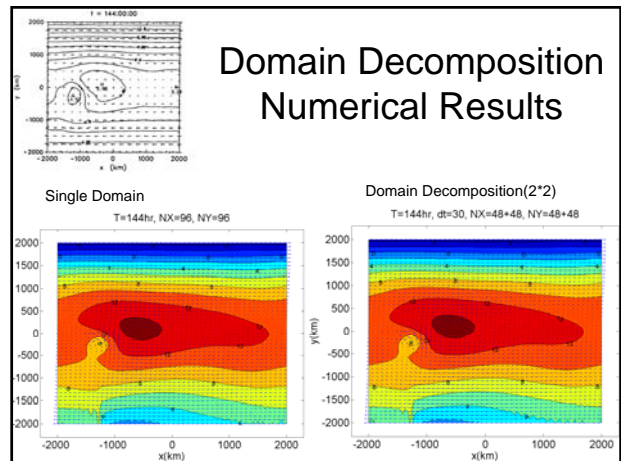
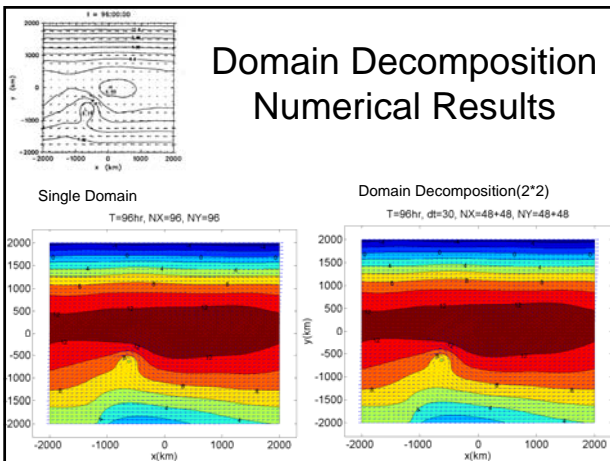
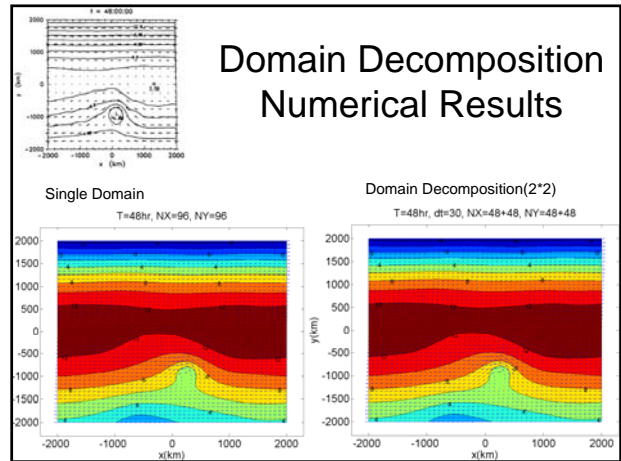
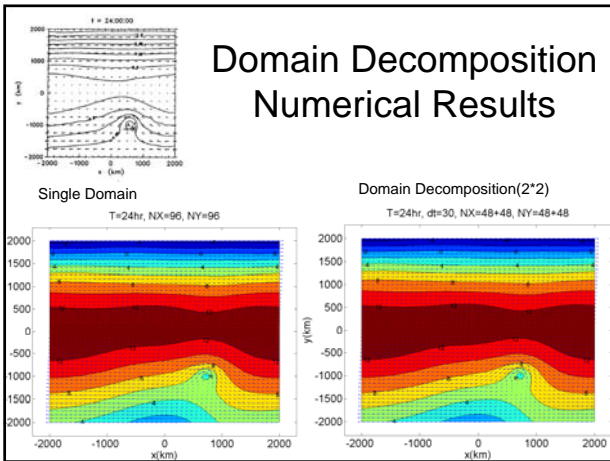
$$\frac{\partial v}{\partial t} + u \frac{\partial v}{\partial x} + v \frac{\partial v}{\partial y} + fu + \frac{\partial h}{\partial y} = 0 \quad (2)$$

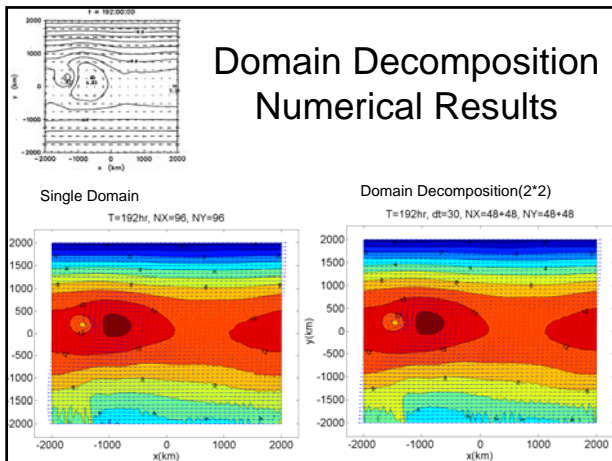
$$\frac{\partial h}{\partial t} + u \frac{\partial h}{\partial x} + v \frac{\partial h}{\partial y} + (\bar{h} + h) \left( \frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} \right) = Q \quad (3)$$

where  $Q(x, y, t) = q_0 \exp\left[-\left(\frac{x-x_0}{\alpha_0}\right)^2 - \left(\frac{y-y_0}{\beta_0}\right)^2\right] \alpha_0^2 \beta_0^2 e^{-t/\tau_0}$

$q_0 = 6250 \text{ m}^2 \text{ s}^{-2}$ ,  $x_0 = y_0 = 200 \text{ km}$ ,  
 $\tau_0 = 6 \text{ hours} = 21600 \text{ sec}$ ,  $(x_0, y_0) = (1000, -1000)$   
 $\bar{h} = c^2 = 2500 \frac{\text{m}^2}{\text{s}^2}$

116

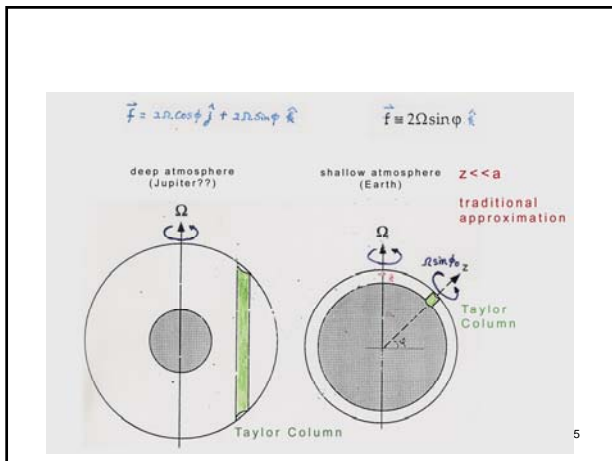
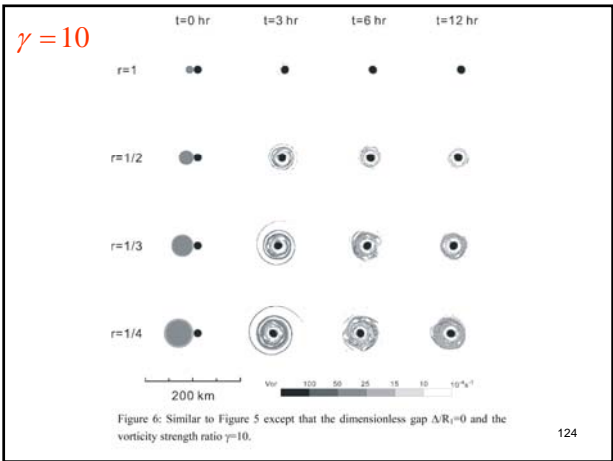
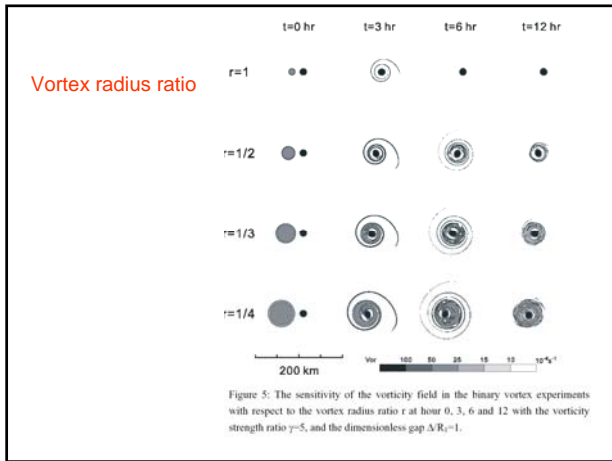




### 問世間颱風是何物？

- 夫大塊噫氣、其名為風 -- 莊子齊物論
- 「颶」-- 明末清初17世紀首見於漢文
- 康熙年間台灣府奏摺：今年亢旱之後、繼以颶風  
今歲入秋亢旱、繼又颶風大作  
台澎海外地方、每直秋潮節候、颶颶時有
- Typhoon -- 英文字根為16世紀阿拉伯語Touffon (旋轉)
- Typhon -- 希臘神話的風神
- Kamikaze -- 神風 (?) 13世紀末蒙古征日

渦旋, 旋轉流體  
Vortex, Rotational Dynamics



### Cascade (Nonlinear Dynamics)

So, nat'ralists observe, a flea  
Hath smaller fleas that on him prey;  
And these have smaller yet to bite 'em,  
And so proceed *ad infinitum*.

----- Jonathan Swift

Big whirls have little whirls  
that feed on their velocity,  
and little whirls have lesser whirls,  
and so on, to viscosity.

----- Lewis Fry Richardson 1922

## Highly Nonlinear System

Turbulence, Order and Chaos  
Multiple scale interaction

**Cascade** of kinetic energy and enstrophy  
Deterministic and statistical dynamics

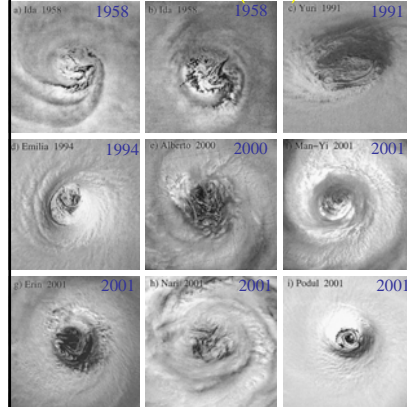
*Laminar yields turbulence*

*Order (i.e. turbulence) emerges from chaos*

**Coherent structures emerge from chaos,**  
*under the action of an external constraint*

127

## Kossin and Schubert (2002)



Vortical Swirls in  
Hurricane Eye Clouds

Eye Dynamics!  
Vortex Mixing!

128

## Symmetric Model

### Ooyama's contribution (1969) — 三層軸對稱模式

Ooyama's experiments have more than 10 years of discussion on sea-air interaction

▲ 「表環流」與「深環流」的交互作用

— “種蛋反應”、“氣旋反應”二尺度的“疊性”

▲ 組織性對流

— 非絕熱效應、積雲參數化的問題

▲ 非線性效應 **linear theory is wrong**

$$\frac{dP}{dt} = \frac{P}{k} Q, \quad \frac{c^2}{f^2} \approx \frac{c^2}{f^2} \left( \frac{f^2 - \partial^2}{f^2} \right) \left( \frac{f^2 - \partial^2}{f^2} \right)$$

▲ 海、氣交互作用  $C_D, C_E$  等皆於

— SST, 初始的 CAPE 不重要

(颱風的維持不能只靠大氣的 CAPE)

▲ 中層的「內流」

— Hurricane spin up

▲ PBL 的乾空氣「進入」

— frictional driven divergence!!!

Cumulus Parameterization

- reduced the multiple scale problems

to one scale problem

(Be careful not to play with

a loaded dice !)

-Buoyancy not exactly zero

Nonlinear Dynamics

Importance of air sea interaction

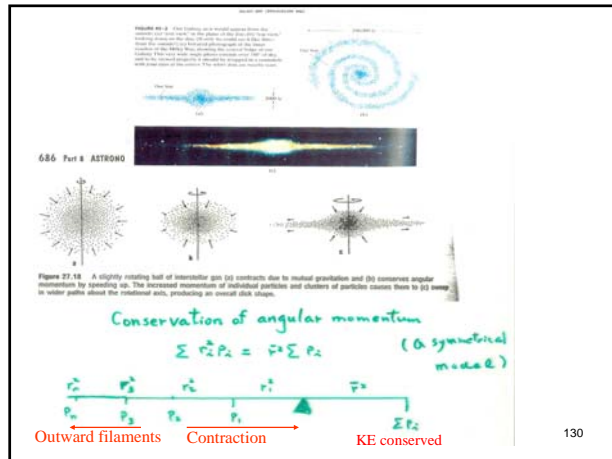
(Initial CAPE NOT important)

Mid-level entrainment

PBL and subsidence

Frictional driven divergence and convergence

129



130

### ● Kirchhoff vortex (nonlinear) Lamb, 1932

$$\frac{x^2}{a^2} + \frac{y^2}{b^2} = 1$$

$a \sim 30 \text{ km}$      $b \sim 20 \text{ km}$

$$\text{angular velocity} = \zeta = \frac{ab}{(a+b)^2} = \omega$$

$$P \sim (2\pi/\zeta) * 4 \quad \zeta \sim 3 * 10^{-3} \text{ s}^{-1}$$

$$\text{rotating period } P = \frac{2\pi (a+b)^2}{\zeta ab}$$

$$V_{\text{max}} \sim 50 \text{ ms}^{-1}$$

### ● Kelvin PV wave (linear)

rotating period P

$$c = V_{\text{max}} \left(1 - \frac{1}{m}\right) \quad m = 2$$

$$\sim \frac{2\pi}{\omega} = \frac{2\pi}{2 \frac{V_{\text{max}}}{r}} * 4 = \frac{2\pi}{\zeta} * 4$$

$$\text{angular velocity} = \frac{c}{r} = \frac{V_{\text{max}}}{2r}$$

Same as Kirchhoff vortex !!

131